

5.15:

Using normalized fields for convenience such that

$$\langle E(t)E(t)^* \rangle = P_1$$

we can write the autocorrelation function in Eq. 5.144 as

$$\langle E(t)E(t-\tau)^* \rangle = P_1 e^{j\omega\tau} e^{-|\tau|/\tau_{coh}} \quad (5.144)$$

Now, if the power in the delayed leg is attenuated by α , such that $P_2 = \alpha P_1$, we can write the power at the detector as

$$\begin{aligned} P_{det} &= \langle [E(t) + \sqrt{\alpha}E(t-\tau)] [E(t) + \sqrt{\alpha}E(t-\tau)]^* \rangle \\ &= \langle E(t)E(t)^* \rangle + \alpha \langle E(t-\tau)E(t-\tau)^* \rangle + \sqrt{\alpha} \langle E(t)E(t-\tau)^* \rangle + \sqrt{\alpha} \langle E(t-\tau)E(t)^* \rangle \\ &= P_1 + \alpha P_1 + \sqrt{\alpha} P_1 e^{j\omega\tau} e^{-|\tau|/\tau_{coh}} + \sqrt{\alpha} P_1 e^{-j\omega\tau} e^{-|\tau|/\tau_{coh}} \\ &= P_1 + P_2 + 2\sqrt{P_1 P_2} \cos(\omega\tau) e^{-|\tau|/\tau_{coh}} \end{aligned}$$

Then, fringes are visible as the value of $\cos(\omega\tau)$ is varied:

$$P_{max} = P_1 + P_2 + 2\sqrt{P_1 P_2} e^{-|\tau|/\tau_{coh}}$$

$$P_{min} = P_1 + P_2 - 2\sqrt{P_1 P_2} e^{-|\tau|/\tau_{coh}}$$

$$P_{max} - P_{min} = 4\sqrt{P_1 P_2} e^{-|\tau|/\tau_{coh}}$$

$$P_{max} + P_{min} = 2(P_1 + P_2)$$

Therefore,

$$e^{-|\tau|/\tau_{coh}} = \frac{P_1 + P_2}{2\sqrt{P_1 P_2}} \frac{P_{max} - P_{min}}{P_{max} + P_{min}} \quad (5.145)$$

The problem asks for the coherence time for the case when $\tau = 3$ ns:

$$e^{-|\tau|/\tau_{coh}} = 0.2$$

$$\tau_{coh} = \frac{-3}{\ln(0.2)} \text{ ns} = 1.864 \text{ ns}$$

The linewidth is then given by Eq. 5.149:

$$\Delta\nu_{FW} = \frac{1}{\pi\tau_{coh}} = 171 \text{ MHz} \quad (5.149)$$

5.16:

From Fig. 5.26, estimate μ_2 near $P_{02} = 2.5$ mW:

$$\mu_2 \equiv \frac{\Delta P_{p-p}}{2P} \approx \frac{0.2 \text{ mW}}{2.5 \text{ mW}} = 0.08$$

Using the effective mirror model with VCSELs, for all practical purposes, we can assume lossless mirrors (The transmission is reduced by $e^{-\alpha L_{eff}}$ according to the discussion surrounding Eq. 3.64. However, $\alpha L_{eff} < 0.005$ represents a negligible reduction in transmission for VCSELs.)

Eq. 5.175 is derived assuming lossless mirrors and can be used to extract κ_f from μ_2 . If all of the light is coupled out of mirror 2, then $F_1 = 0$ and since $F_1 + F_2 = 1$ for lossless mirrors, we have shown that $F_2 = 1$. The problem also says to assume that we are well above threshold; so we can neglect the third term in defining μ_2 . With $F_2 = 1$, we can write Eq. 5.175 as:

$$\mu_2 = 2\kappa_f \tau_p \left(\frac{\eta_i}{\eta_d} - 1 \right)$$

Solve for κ_f :

$$\begin{aligned} \kappa_f &= \frac{\mu_2}{2\tau_p \left(\frac{\eta_i}{\eta_d} - 1 \right)} \\ &= \frac{0.08}{2(2.77 \text{ ps}) \left(\frac{0.8}{0.494} - 1 \right)} ; \tau_p = 2.2 \text{ ops} \\ &= \frac{2.331 \times 10^{10} \text{ s}^{-1}}{2.935} \end{aligned}$$

$$\kappa_f = \frac{t_2^2 \sqrt{f_{ext}}}{r_2 \tau_L}$$

Solve for f_{ext} :

$$f_{ext} = \left(\frac{\kappa_f r_2 \tau_L}{t_2^2} \right)^2$$

From Table 5.1,

$$r_2 = \sqrt{R_2} = 0.995$$

$$r_1 = 1$$

$$t_2^2 = 1 - 0.995^2 = 0.00998 \text{ assuming lossless mirrors}$$

$$\tau_L = \frac{2L}{v_g} = \frac{2(10^{-4} \text{ cm})}{4.2 \times 10^{10} \text{ cm/s}} = 2.8 \times 10^{-14} \text{ s}$$

$$\begin{aligned} f_{ext} &= \left(\frac{(2.331 \times 10^{10} \text{ s}^{-1})(0.995)(2.8 \times 10^{-14} \text{ s})}{0.00998} \right)^2 \\ &= \frac{6.706}{4.23} \times 10^{-3} \end{aligned}$$

$$\begin{aligned} f_{ext}(\text{dB}) &= 10 \log(f_{ext}) \\ &= -23.7 \text{ dB} \\ &= -21.73 \text{ dB} \end{aligned}$$



5.17:

From Fig. 5.26 and Eq. 5.175, $\cos(2\beta L_p)$ goes through one complete period for a change in current of 0.4 mA. Assuming that the change of β with current is approximately linear, we can write

$$\frac{d(2\beta L_p)}{dI} = \pm \frac{2\pi}{0.4 \text{ mA}}$$

$$\beta \equiv \frac{2\pi\bar{n}}{\lambda_0}$$

Neglecting the change in index due to current and temperature, we can write the first equation above as

$$2L_p(2\pi\bar{n})\frac{d\frac{1}{\lambda_0}}{dI} = \pm \frac{2\pi}{0.4 \text{ mA}}$$

which can be rearranged:

$$L_p = \pm \frac{1}{(0.4 \text{ mA}) \bar{n} 2 \frac{d(\frac{1}{\lambda_0})}{dI}}$$

and simplified

$$L_p = \pm \frac{-\lambda_0^2}{(0.8 \text{ mA})\bar{n} \left(\frac{d\lambda_0}{dI}\right)}$$

All variables are known except $\frac{d\lambda_0}{dI}$. Assuming that the wavelength change with current is due solely to the change in temperature, we can write

$$\frac{d\lambda_0}{dI} = \frac{d\lambda_0}{dT} \frac{dT}{dP_d} \frac{dP_d}{dI}$$

where T and P_d denote temperature and dissipated power, respectively.

$$\frac{d\lambda_0}{dT} = 0.08 \frac{\text{nm}}{^\circ\text{C}}$$

given in problem

$$\frac{dT}{dP_d} = 3 \frac{^\circ\text{C}}{\text{mW}}$$

given in problem

$$P_d = P_{in} - P_{out} = IV - P_0$$

$$\frac{dP_d}{dI} = \frac{dP_{in}}{dI} - \frac{dP_{out}}{dI} = 3V - \frac{2.5 \text{ mW} - 0 \text{ mW}}{6.5 \text{ mA} - 1.75 \text{ mA}} = 2.47 \frac{\text{mW}}{\text{mA}}$$

$$\frac{d\lambda_0}{dI} = 0.59 \frac{\text{nm}}{\text{mA}}$$

Then, plugging back into the equation for L_p , we have

$$L_p = \frac{(980 \text{ nm})^2}{(0.8 \text{ mA})(4.2) \left(0.59 \frac{\text{nm}}{\text{mA}}\right)}$$

$$L_p = \text{substrate thickness} = 480 \mu\text{m}$$

5.18:

a)

$$\kappa_f = \frac{v_g t_2^2 \sqrt{f_{ext}}}{2Lr_2} \quad (\text{p. 247})$$

$$\kappa_f = \frac{v_g(1-r_2^2)(r_3 t_c)}{2Lr_2}$$

$\bar{n}_g$ is not specified by the problem. Assume that laser is made in the InGaAsP material system and the mode has $\bar{n}_g = 4.0$ as mentioned on p. 40.

Assume a facet reflectivity $r_2^2 = 0.32$.

$$\kappa_f = \frac{\frac{3.0 \times 10^{10} \text{ cm/s}}{4.0} (1 - 0.32) (\sqrt{0.04} \sqrt{0.25})}{2(0.03 \text{ cm}) \sqrt{0.32}} = 1.50 \times 10^{10} \text{ s}^{-1}$$

$$\tau_{ext} = \frac{2n_{fiber}L_p}{c} = \frac{2(1.45)L_p}{3 \times 10^{10} \text{ cm/s}} = (0.967 \times 10^{-5} \text{ s/km})L_p \quad (\text{p. 250})$$

$$C = \kappa_f \tau_{ext} \sqrt{1 + \alpha^2} = (1.45 \times 10^5 \text{ km}^{-1})L_p \sqrt{1 + \alpha^2} \quad (5.180)$$

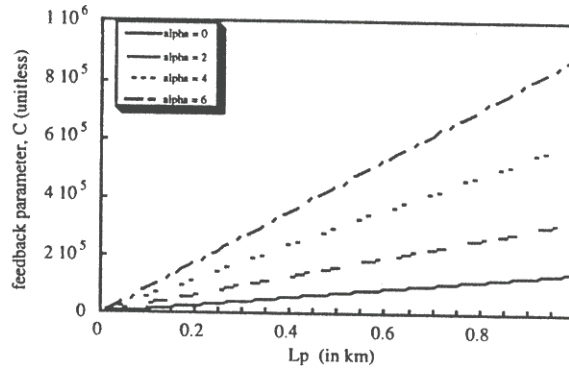


Figure 5.18a. Feedback coefficient, C , vs. fiber length, L_p .

b)

$$\frac{\Delta\nu}{\Delta\nu_0} = \frac{1}{(1 + C \cos(2\beta L_p + \phi_\alpha))^2} \quad (5.181)$$

Since $\cos(2\beta L_p + \phi_\alpha)$ can be adjusted to any value between -1 and +1, we can break this equation into 2 cases:

Case 1: $C < 1$:

$$\text{minimum } \frac{\Delta\nu}{\Delta\nu_0} = \frac{1}{(1+C)^2}$$

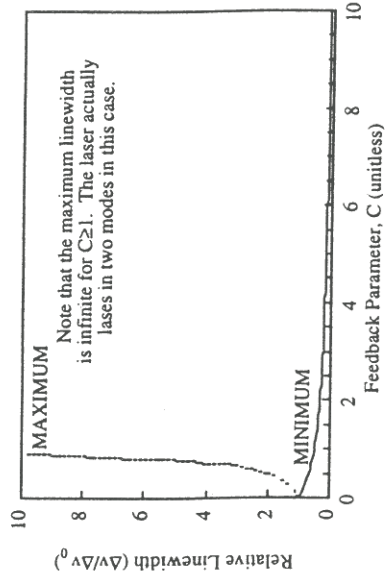
$$\text{maximum } \frac{\Delta\nu}{\Delta\nu_0} = \frac{1}{(1-C)^2}$$

Case 2: $C \geq 1$:

$$\text{minimum } \frac{\Delta\nu}{\Delta\nu_0} = \frac{1}{(1+C)^2}$$

$$\text{maximum } \frac{\Delta\nu}{\Delta\nu_0} = \infty$$

For $C \geq 1$, $\cos(2\beta L_p + \phi_\alpha)$ can be adjusted to be equal to $1/C$, which yields an infinite bandwidth. In this case, the laser has split into two lasing modes, which implies that $\Delta\nu$ is meaningless for this case.

Figure 5.18b. Laser linewidth vs. feedback coefficient, C .