

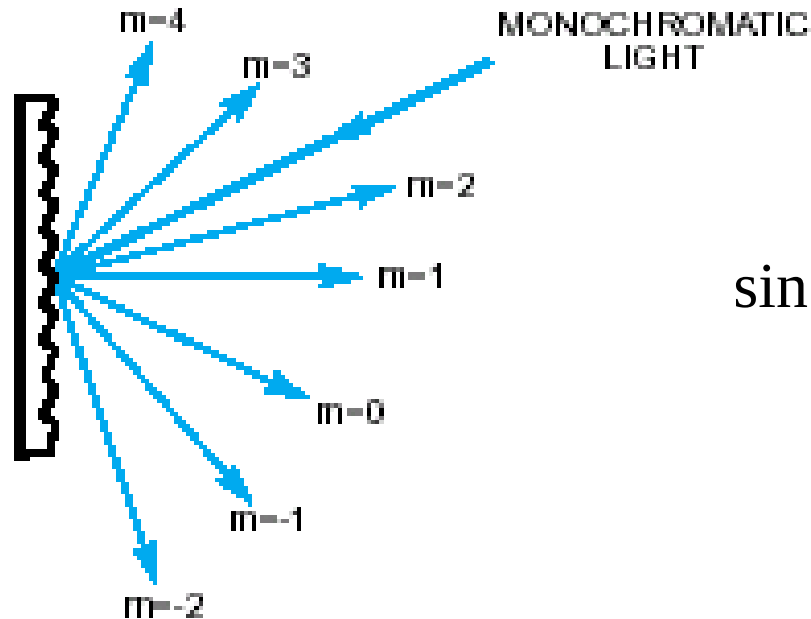
Spectroscopy

- Diffraction gratings
- Resolution
- Throughput
- Interferometers
- Filters

Terminology

- Spectrometers measure wavelengths, monochromators separate wavelengths
 - Scanning grating
 - Fixed grating with linear array
 - Fixed grating with area array
- Laser spectroscopy: tunable laser as monochromator
- Interferometers
- Filters

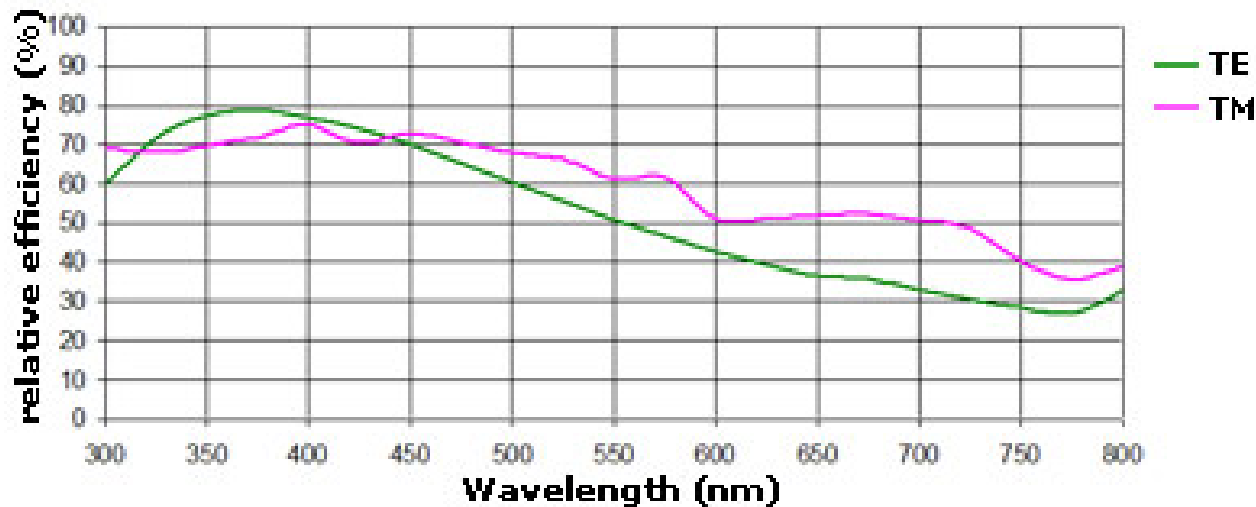
Diffraction Gratings



$$\sin \theta_i + \sin \theta_d = m \frac{\lambda}{a}$$

- Ruled gratings: up to 1200 lines/mm ($a = 0.83 \mu\text{m}$)
- Holographic gratings: 150-5670 lines/mm
- Grating width, W: 1 mm – 12 cm

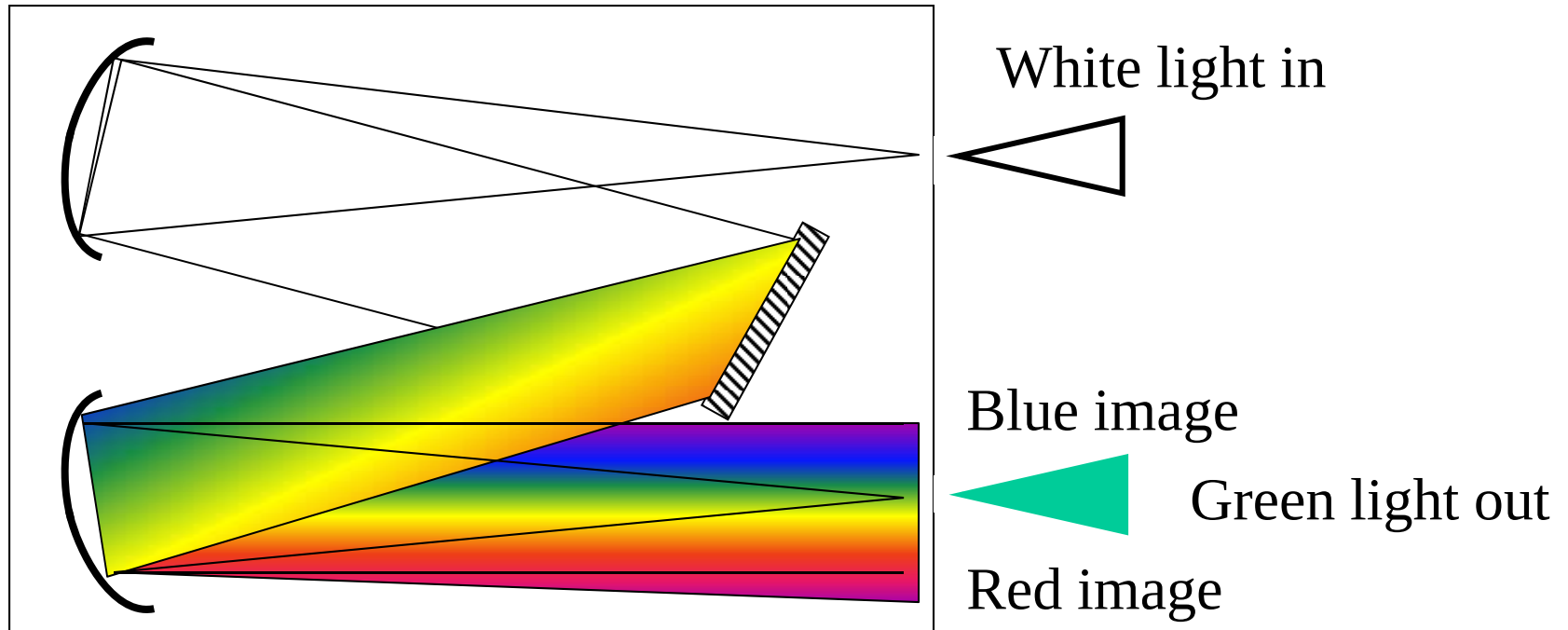
Grating Efficiency



510 12; 600 l/mm; blaze 400 nm

- Polarization dependent
- Blazed gratings shift energy into particular order
- Efficiency falls fast at $\lambda < \lambda_{\text{blaze}}$

Grating Monochromator



Mirrors determine system focal length f ,
and NA ($f/\#$)

Resolution

Linear Dispersion: $\frac{dL}{d\lambda} = \frac{fm}{a \cos \theta_d}$

Slit-limited resolution: $\Delta\lambda = \frac{\text{slit width}}{\text{linear dispersion}}$

Pixel-limited resolution: $\Delta\lambda = \frac{1 - 3 \text{ times pixel width}}{\text{linear dispersion}}$

Resolving Power: $R = \frac{\lambda}{\Delta\lambda} = \frac{W(\sin \theta_i + \sin \theta_d)}{\lambda}$

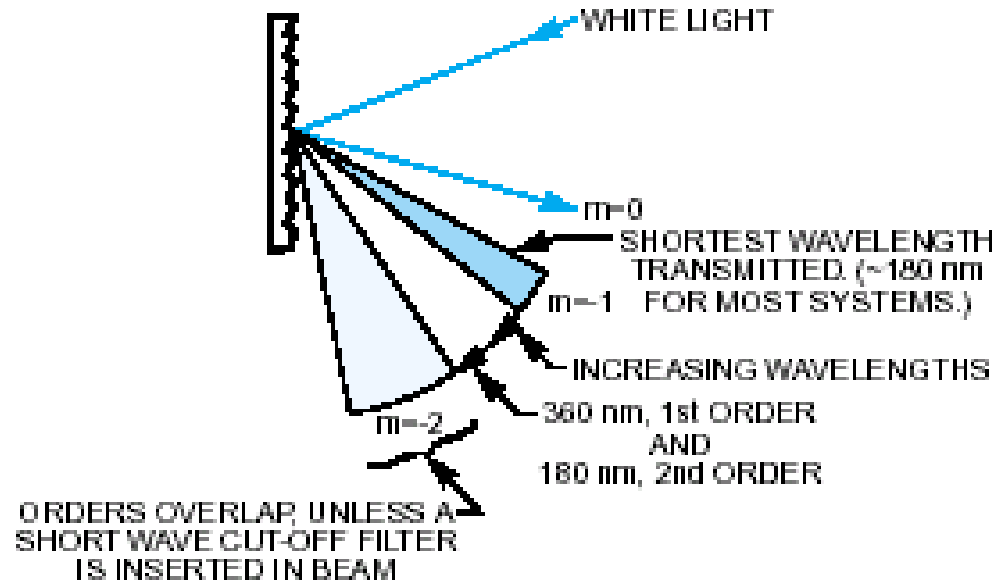
Aberration-limited resolution: coma $\sim \text{NA}^3$

Throughput

$$\frac{P_{\text{out}}}{P_{\text{in}}} = \text{Grating Efficiency} \cdot \text{Reflection losses} \cdot \left[\frac{NA_{\text{mono.}}}{NA_{\text{illum.}}} \right]^2$$

- A monochromator is an optical system:
consider geometrical extent
- Don't *dramatically* decrease NA_{illum} to
increase throughput: resolution will suffer

Free Spectral Range

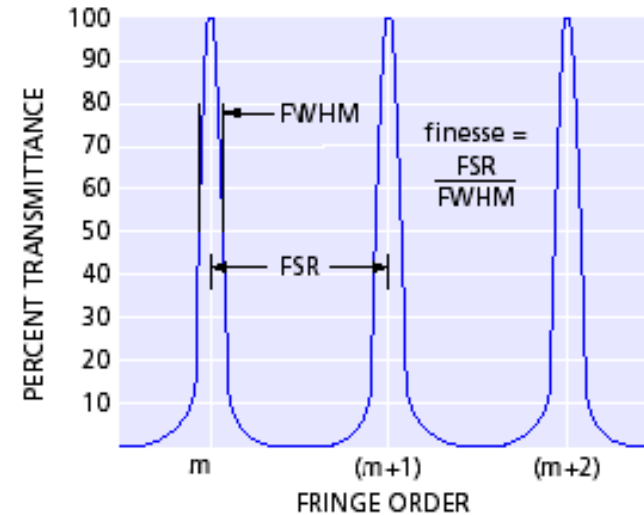
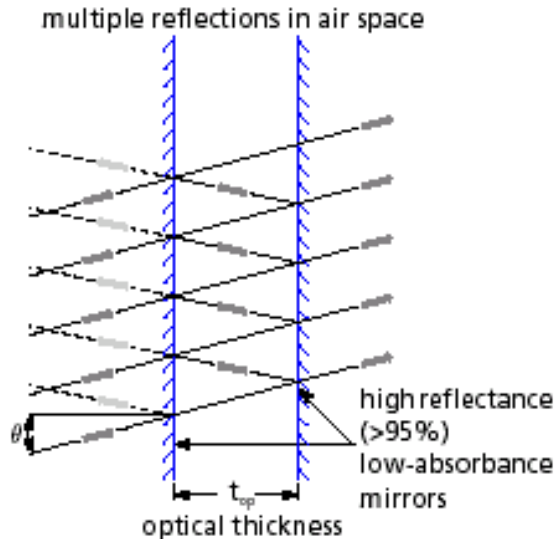


- Spectra from different diffraction orders can overlap
- Use optical filters to narrow spectral range to single order
- Order 0 has no dispersion: useful for alignment

Interferometers

- Fabry Perot
 - tunable filters
 - high resolution optical spectrum analyzers
 - interference filters
- Michelson
 - FTIR spectrometers
 - Wavemeters
- Mach-Zehnder
 - Fiber interferometers
 - Integrated optics

Fabry Perot Interferometer

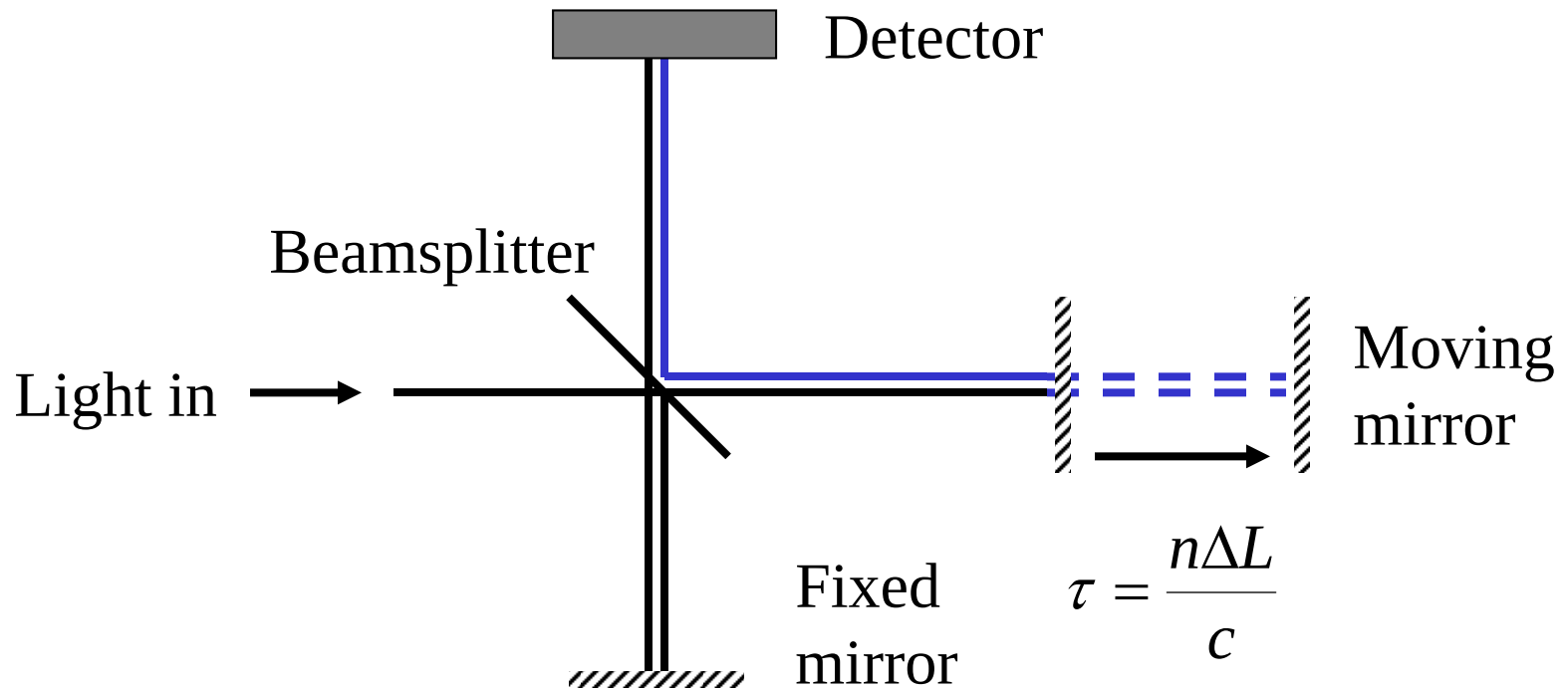


$$\Delta\lambda_{1/2} = \frac{\lambda^2}{2\pi n l} \frac{1-R}{\sqrt{R}}$$

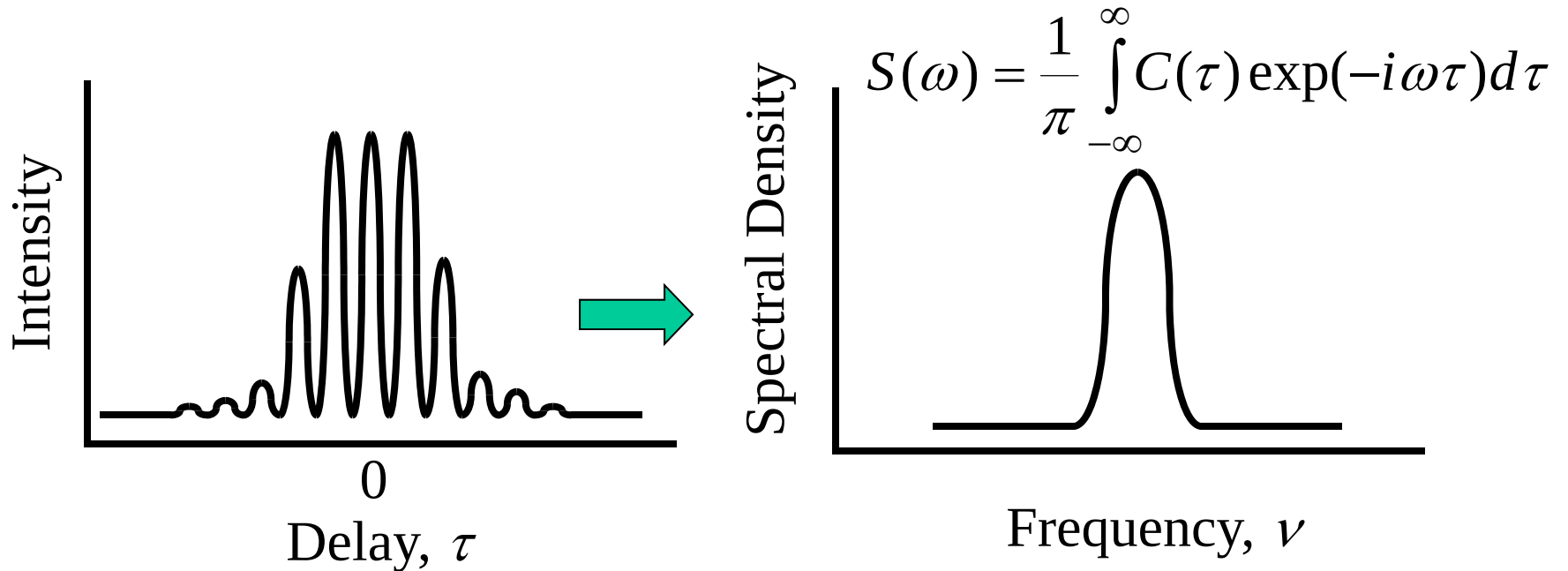
$$FSR = \frac{\lambda^2}{2nl}$$

- Vary gap width to scan resonant wavelength
- Cascade 2nd FPI or grating monochromator for higher FSR

Michelson Interferometer



Fourier Transform Spectroscopy

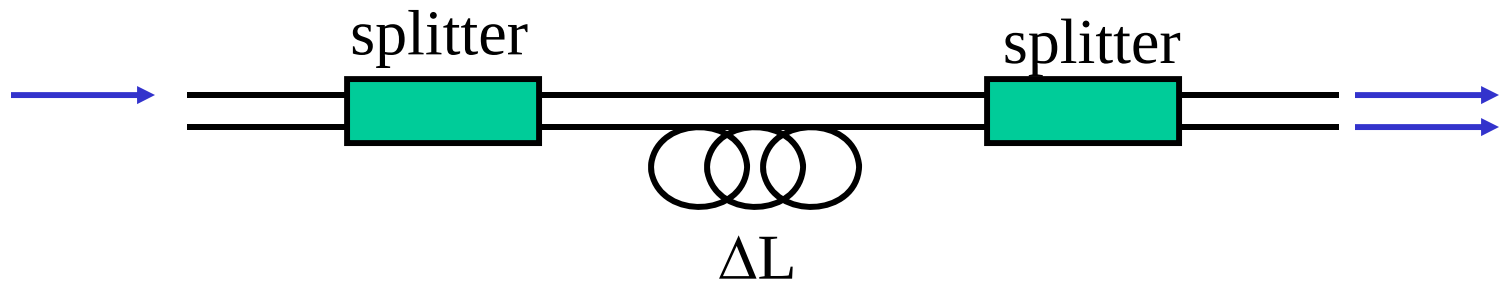


Throughput advantages:

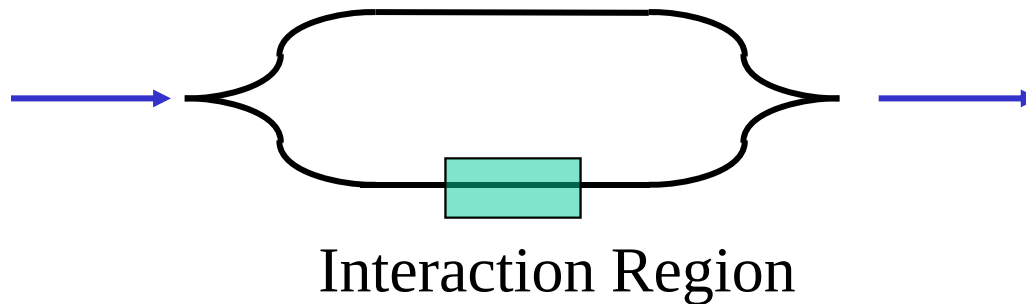
- Large collecting aperture (no slits)
- All wavelengths measured at the same time

Mach Zehnder Interferometer

Fiber Implementation: Very long ΔL possible



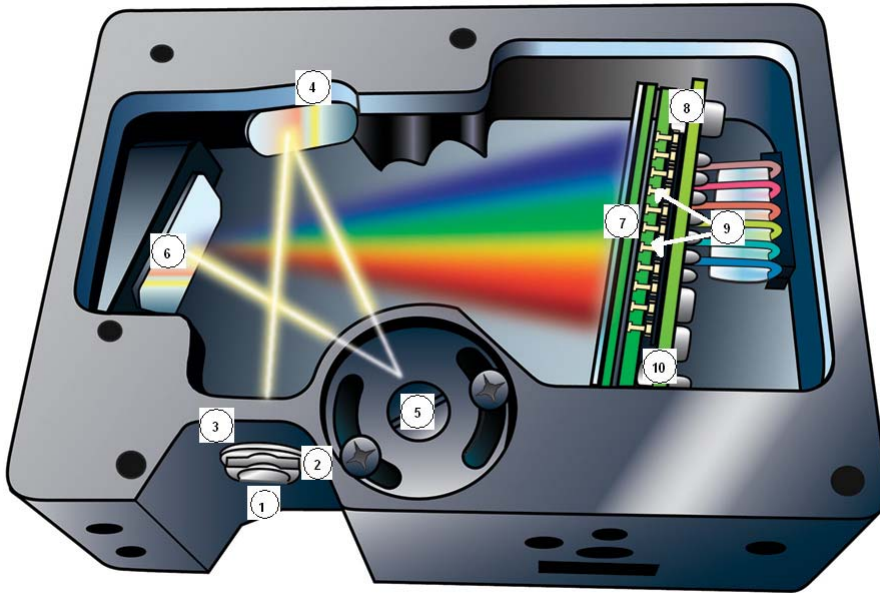
Planar waveguide implementation: Mass-producible



Optical Spectrum Analyzers

- Fixed grating spectrometers with CCDs or Diode Arrays
- Scanning grating spectrometers
- Interferometers

Miniature Computerized Spectrometers



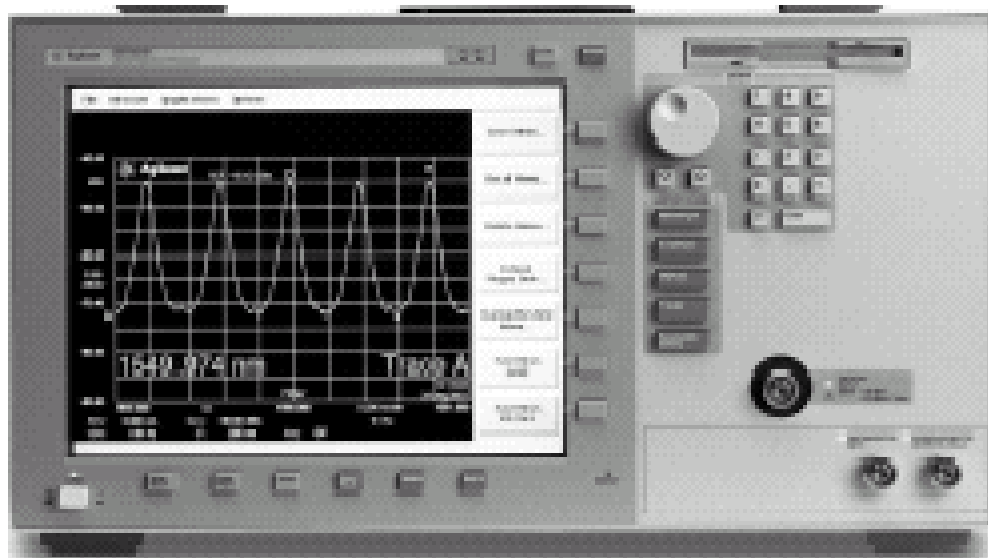
- Palm sized
- Si CCD or InGaAs array
- ~ 1 nm resolution
- USB driven
- Fiber or free space coupling
- Fast and low cost

Spectrometer and Diode Array



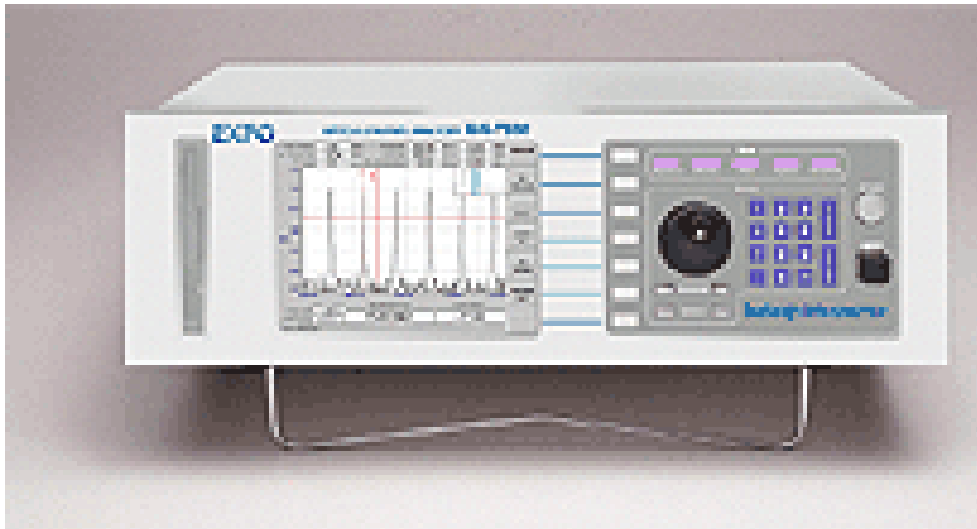
- Flexible configuration
- Free space or fiber coupled
- Components controlled by external computer
- Not portable

Optical Spectrum Analyzer



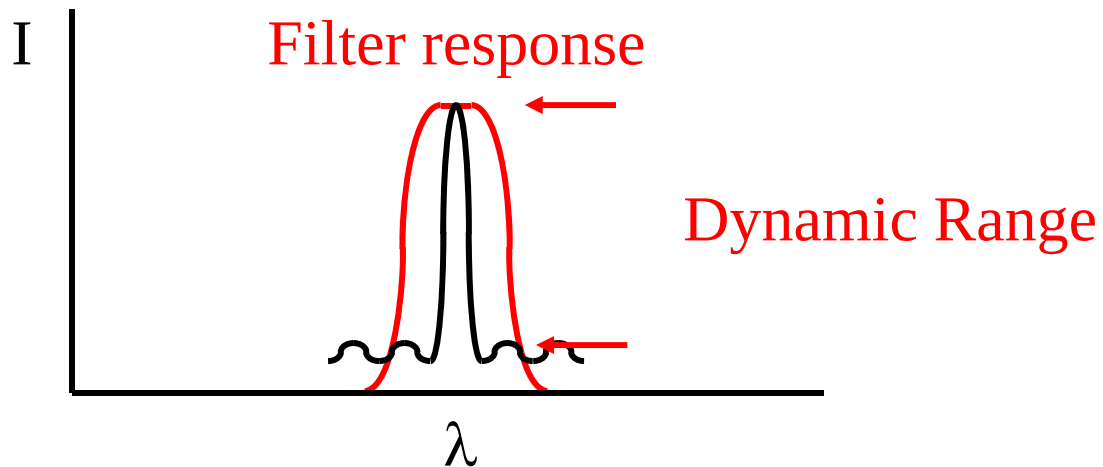
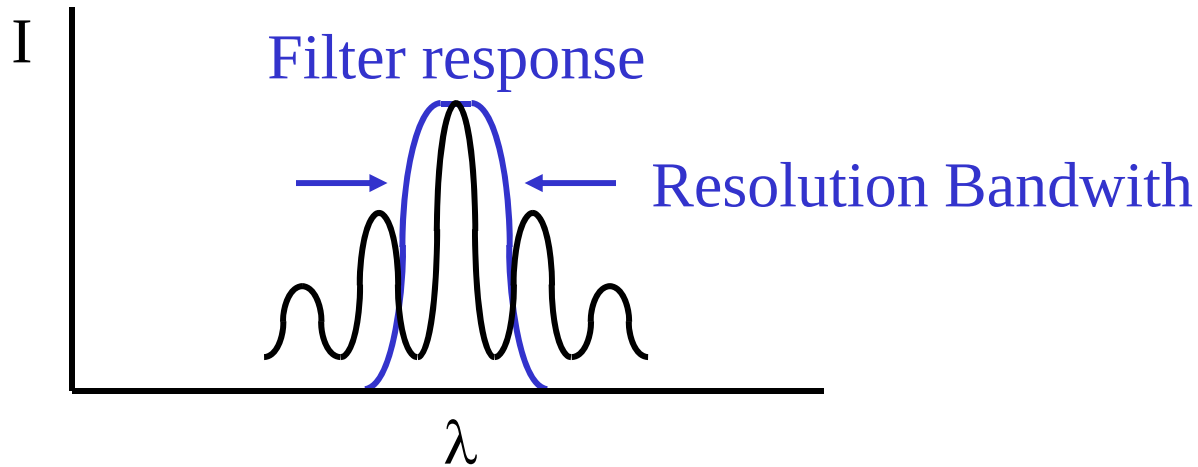
- Fiber input
- Computerized
- ~ 0.05 nm resolution

Wavemeter



- Interferometer based
- Fiber input
- Computerized

OSA Definitions



dB and dBm

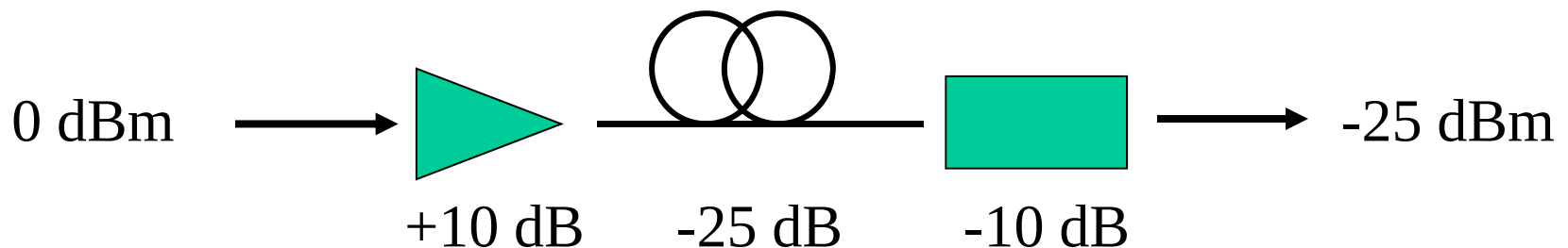
Decibels(dB) refer to a POWER ratio:

$$\text{dB} = 10\text{LOG}(P_{\text{out}}/P_{\text{in}})$$

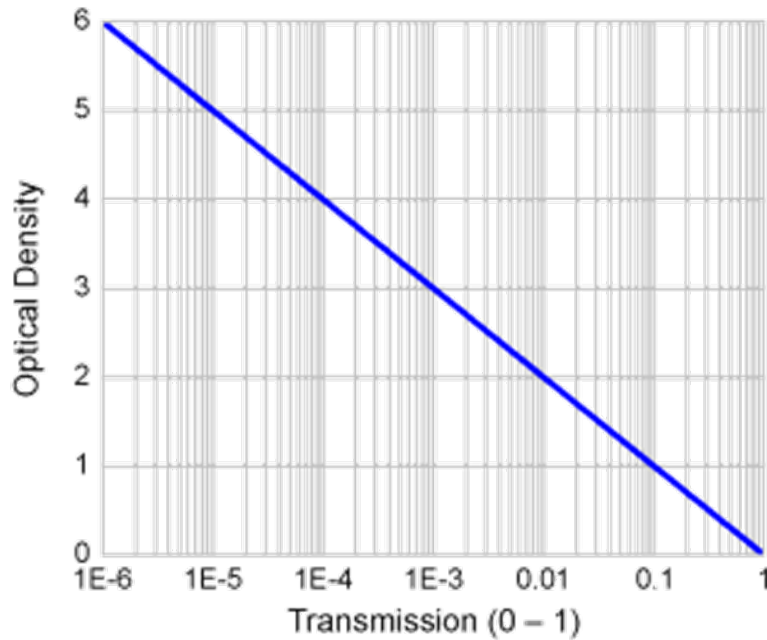
dBm measures relative to 1 mW

$$\text{dBm} = 10\text{LOG}(P_{\text{out}}/1 \text{ mW})$$

Losses through system add, if measured in dB.



Neutral Density Filters



- “Neutral” = wavelength independent
- Optical Density $OD = -\log_{10}(T)$

Interference Filters

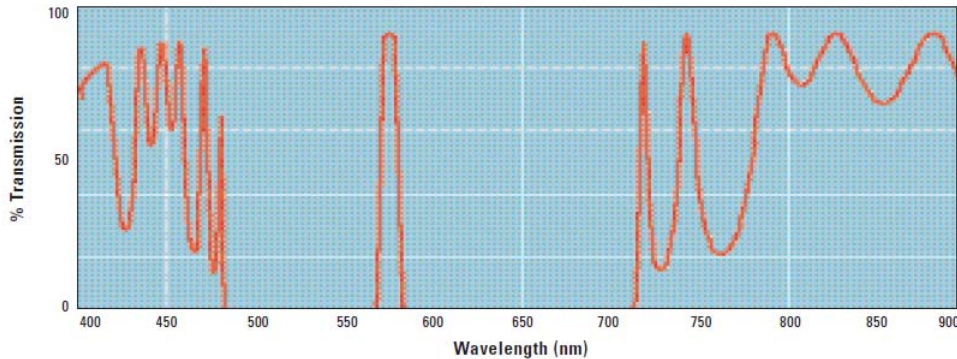


Figure 3: Bandpass Without Extended Attenuation

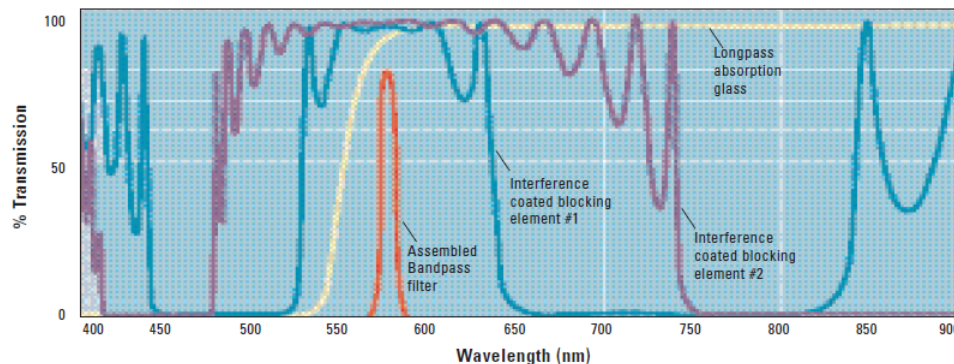


Figure 4: Bandpass With Extended Attenuation

- Fabry Perot Cavity
- Limited blocking beyond resonance
- Extended blocking with additional layers, absorbers
- Blocking layers reduce T, and can fluoresce

Angle and Polarization Response

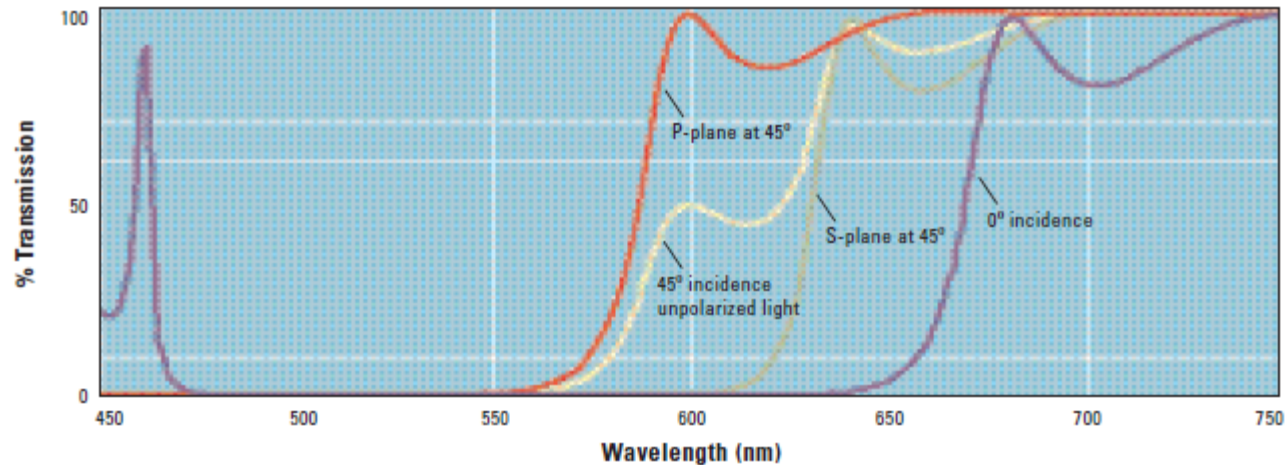


Figure 5: Angle of Incidence Polarization Effects—Longpass Filter

- Increasing angle of incidence shifts edges to shorter wavelength
- Filters are polarization dependent at non-normal incidence

Credits

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- www.omegafilters.com
- Yariv, A., Optical Electronics
- Hecht and Zajac, Optics